



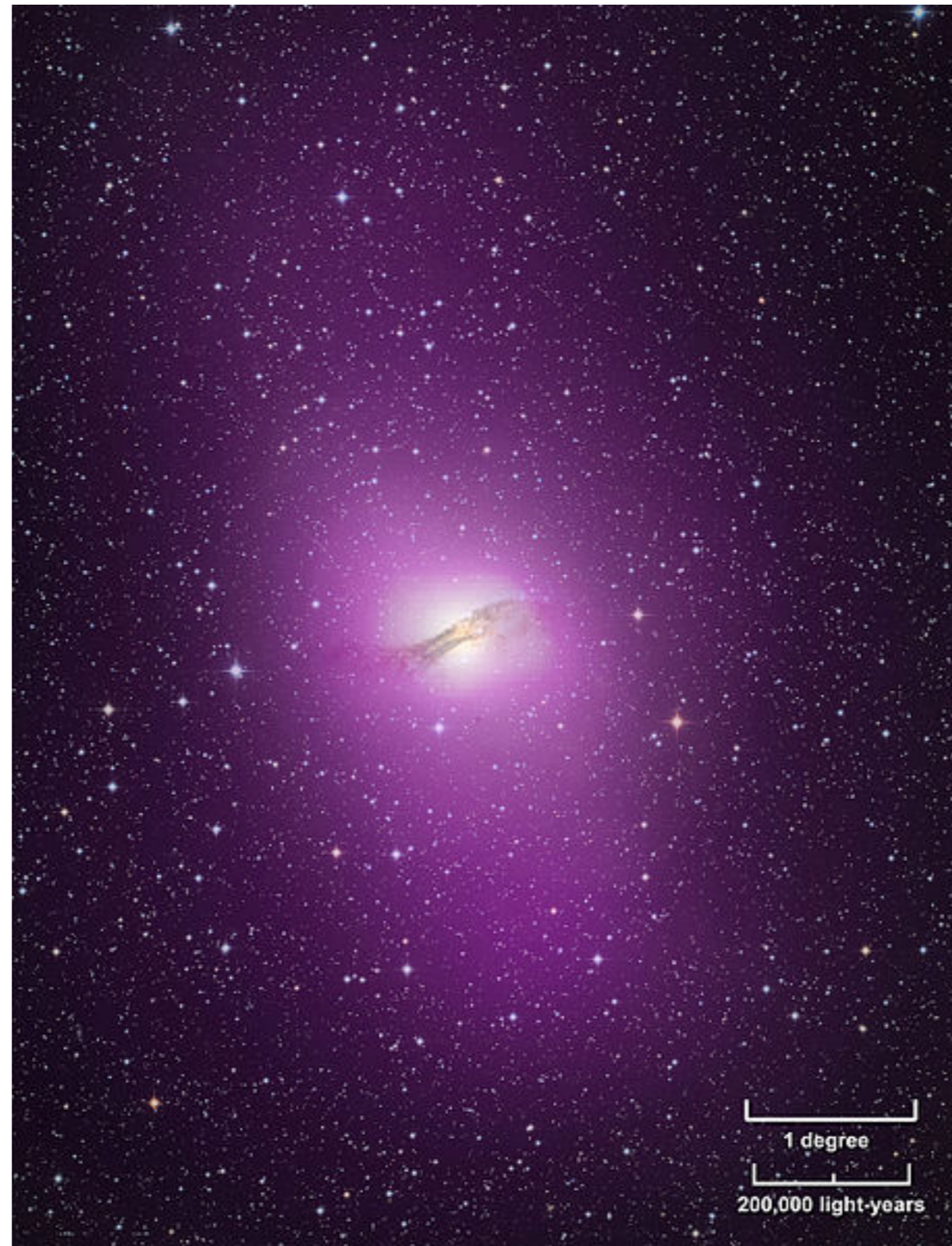
AST 1420

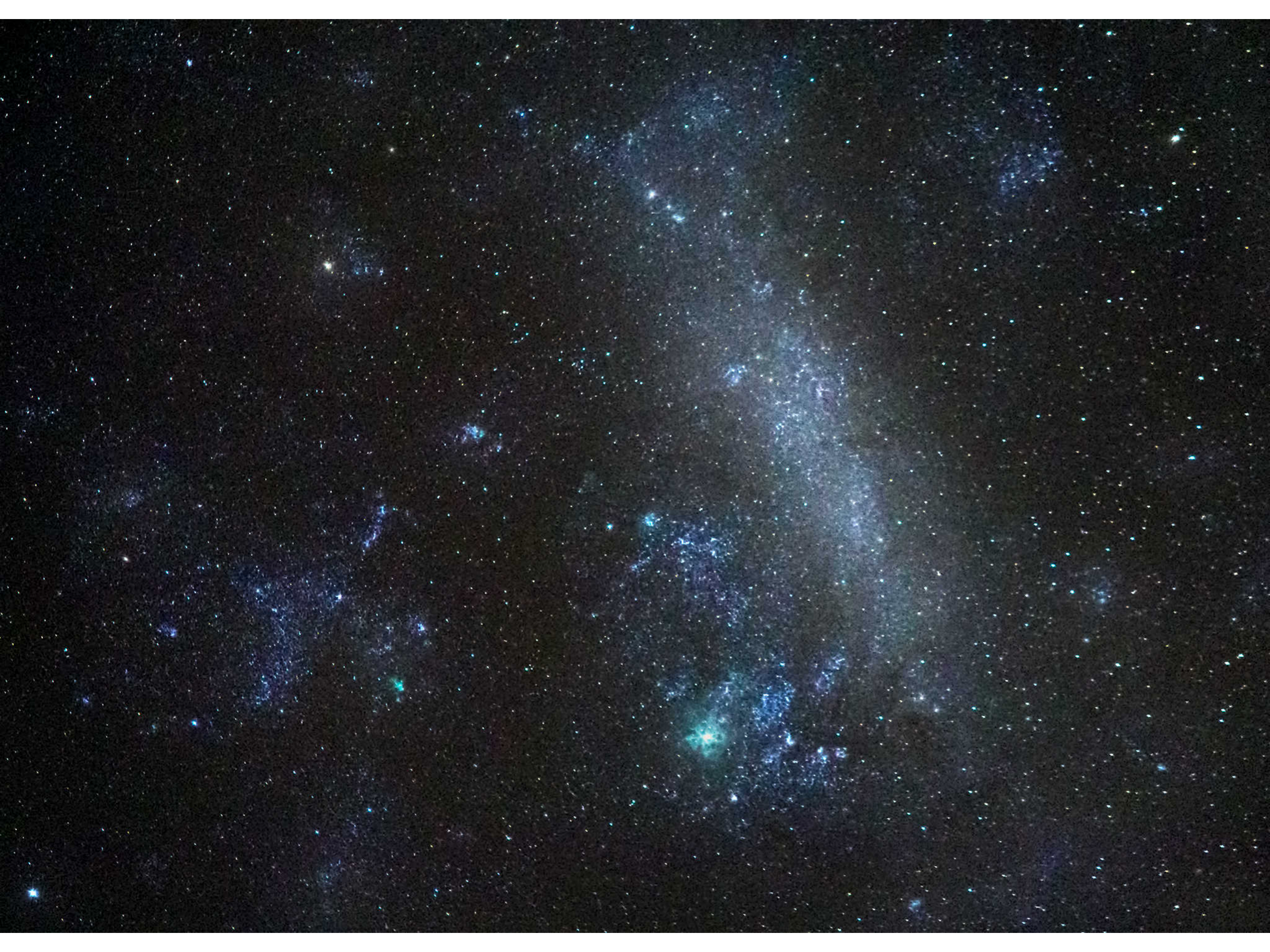
Galactic Structure and Dynamics

M51



Cen A







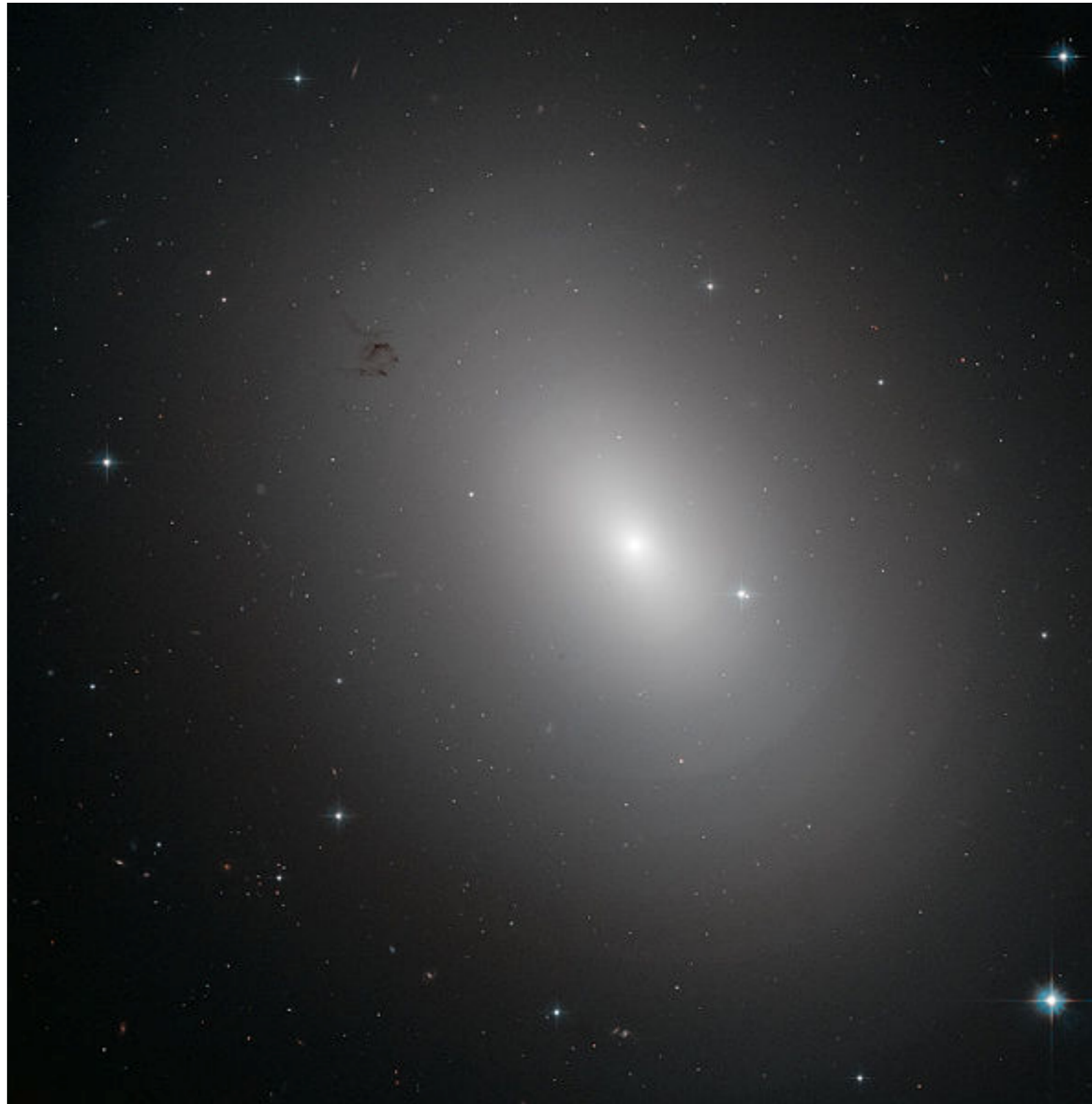
NGC 1300





M81

NGC 3923





0.0 Gyr

Credit: Greg Stinson, MUGS (<http://mugs.mcmaster.ca/>)

# Why study galaxies?

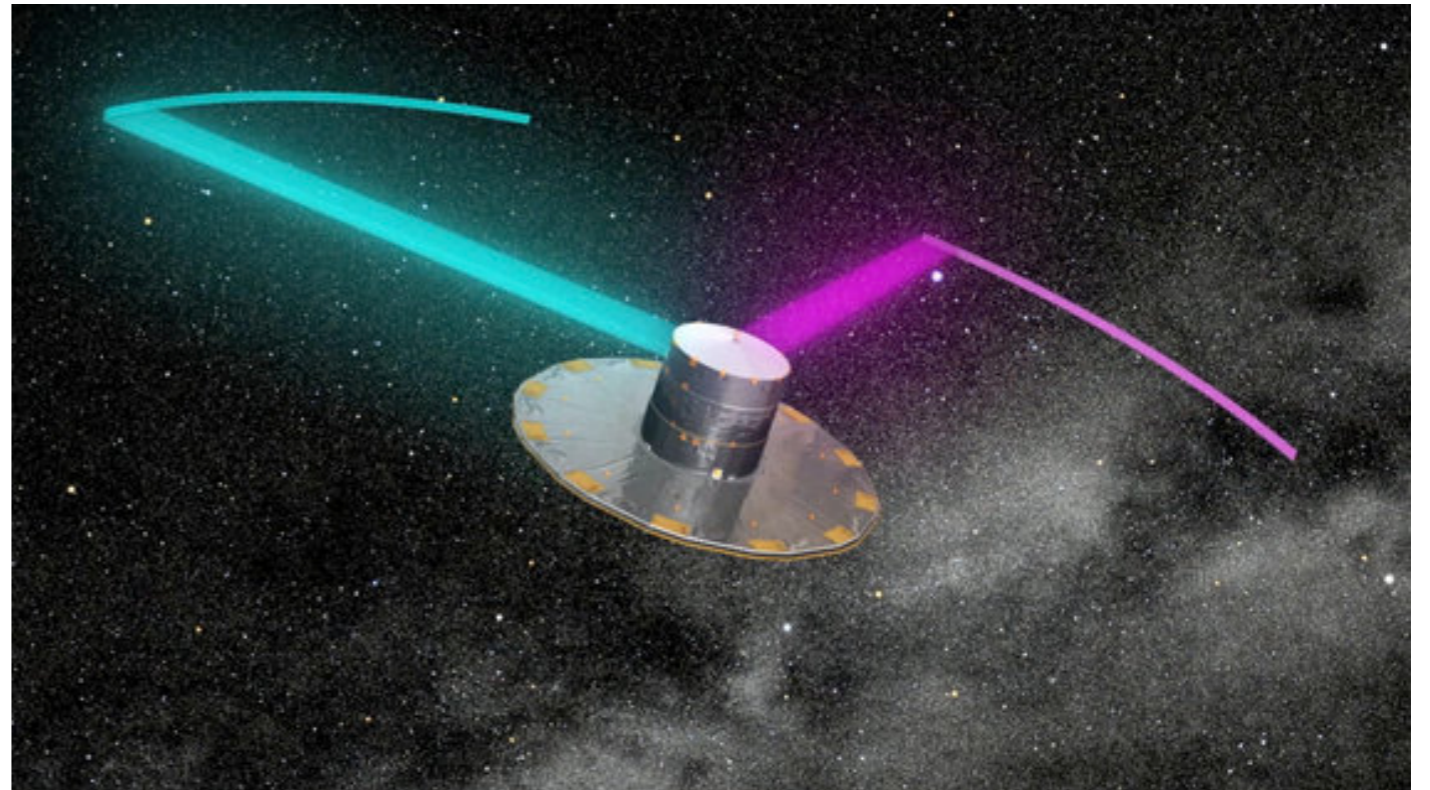
- Fascinating cosmic objects!
- Great application of fundamental physics: GR: galaxy formation in expanding Universe; Newtonian gravity dominating the evolution of bound galaxies; radiation, hydrodynamics, magnetic fields,...
- Our own cosmic genesis: how did the Milky Way that contain our solar system form? Where did the solar system travel over the lifetime of the Sun?
- Cosmic laboratories for investigating *dark matter*

# Why study Galactic Structure and Dynamics?

- Gravity is the dominant force in galaxies: most of the mass *only*\* feels gravity (stars and dark matter)
- Could just run large simulations but:
  - Running large, gravity-only simulations still *very* expensive, don't always lead to a very good *understanding* of gravitational effects
  - Additional physics (“baryonic physics”) of star-formation, feedback from stellar winds, supernovae, active galactic nuclei very uncertain *and* difficult to simulate
- Newtonian gravity + dark matter: simple framework to understand complex phenomenology of galaxies
- Only well-understood physical systems can lead to *big discoveries*: e.g., dark matter, dark energy

# Golden age of galactic dynamics

- *Gaia* satellite is scanning the sky and making high-precision measurements of stellar positions over five years —> measure stellar distances, motions, and stellar properties for >1 billion stars!



- First major data release in April 2018, next one on Dec. 3 2020!
- Is providing an incredibly detailed view of all aspects of galactic dynamics: detailed kinematics in the disk, most precise measurement of structure of dark matter halo of *any* galaxy, internal kinematics of clusters, star-forming regions, globular clusters, orbits of all satellite galaxies, ...

# Objectives of this course

- To know and understand the basic physical properties of galaxies: constituents of galaxies, their dynamics, and relation to each other
- Overview of types of tools available for studying galaxy formation and evolution
- Hone astrophysical problem solving skills: combination of analytical thinking, numerical approaches, simulations, and data analysis

# Course details

- Full details on the website:  
<https://github.com/jobovy/AST1420>
- Lectures:
  - Asynchronous, posted on MS Stream, linked on the course website
  - Weekly Q&A session: **Thu at 12:10 pm** (to 1 or 1:30 pm)
- Email: [jo.bovy@utoronto.ca](mailto:jo.bovy@utoronto.ca)
- Getting help:
  - Slack channel #ast1420-fall2020 (you may want to set up a channel on the GASA slack, make sure to invite non-astro students)
  - Email to set up a Zoom appointment



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# AST1420

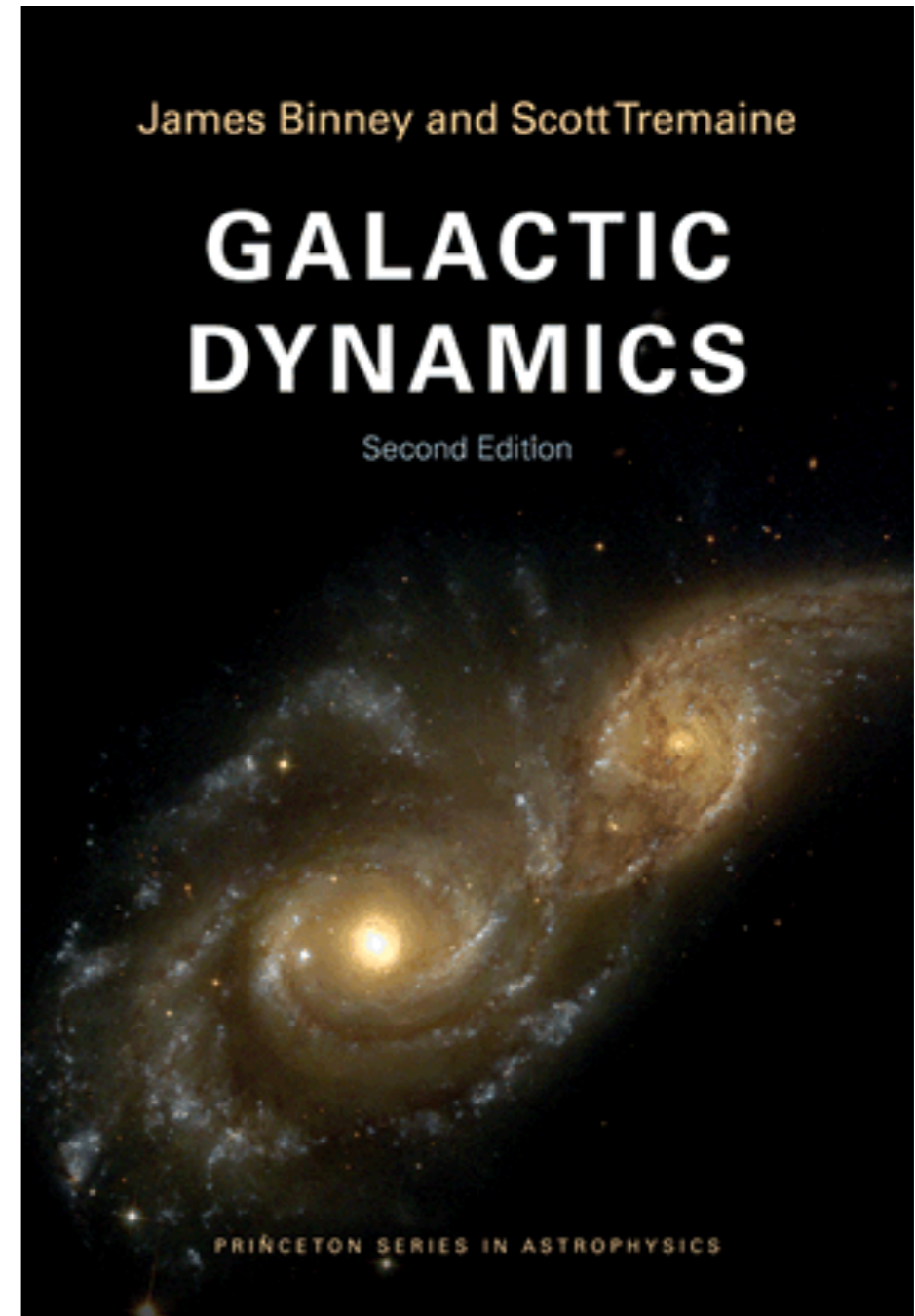
# Lecture notes

- Linked to from course webpage
- New notes will be posted at least one week ahead of class (but most already there; may get slightly updated though)
- Some webpages have lots of content / math-to-typeset; you might want to keep these pages open in different tabs
- Should work in Chrome, may have issues in other browsers
- Please send me an email pointing out any typos, broken links, or other mistakes in the notes



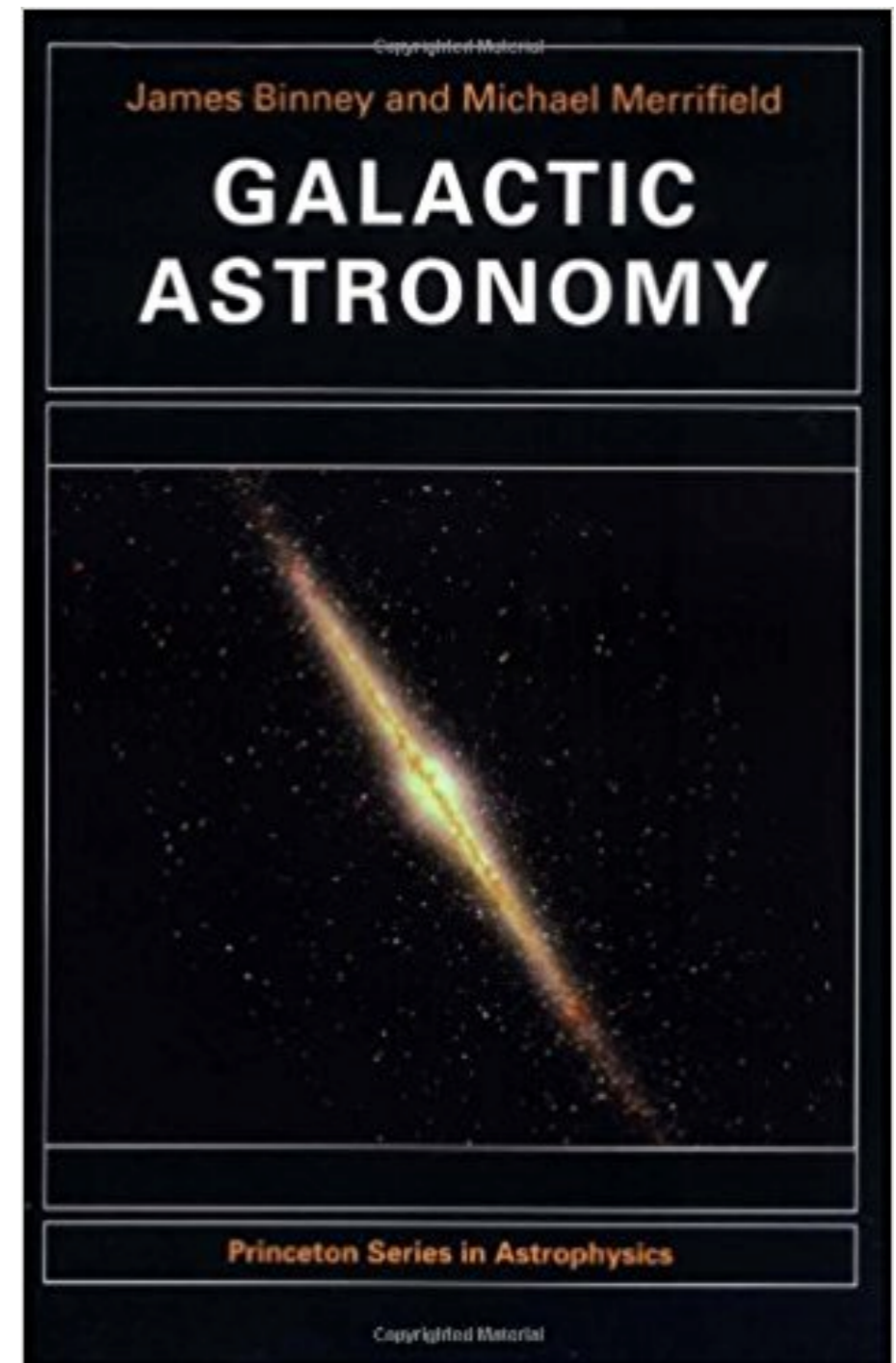
# Additional reading

- Essential reference book:  
Binney & Tremaine, *Galactic Dynamics, 2nd Edition, 2008*,  
Princeton University Press
- Goes into more detail on  
some topics than the notes  
will + advanced material
- Must-have for the galactic  
dynamicist!



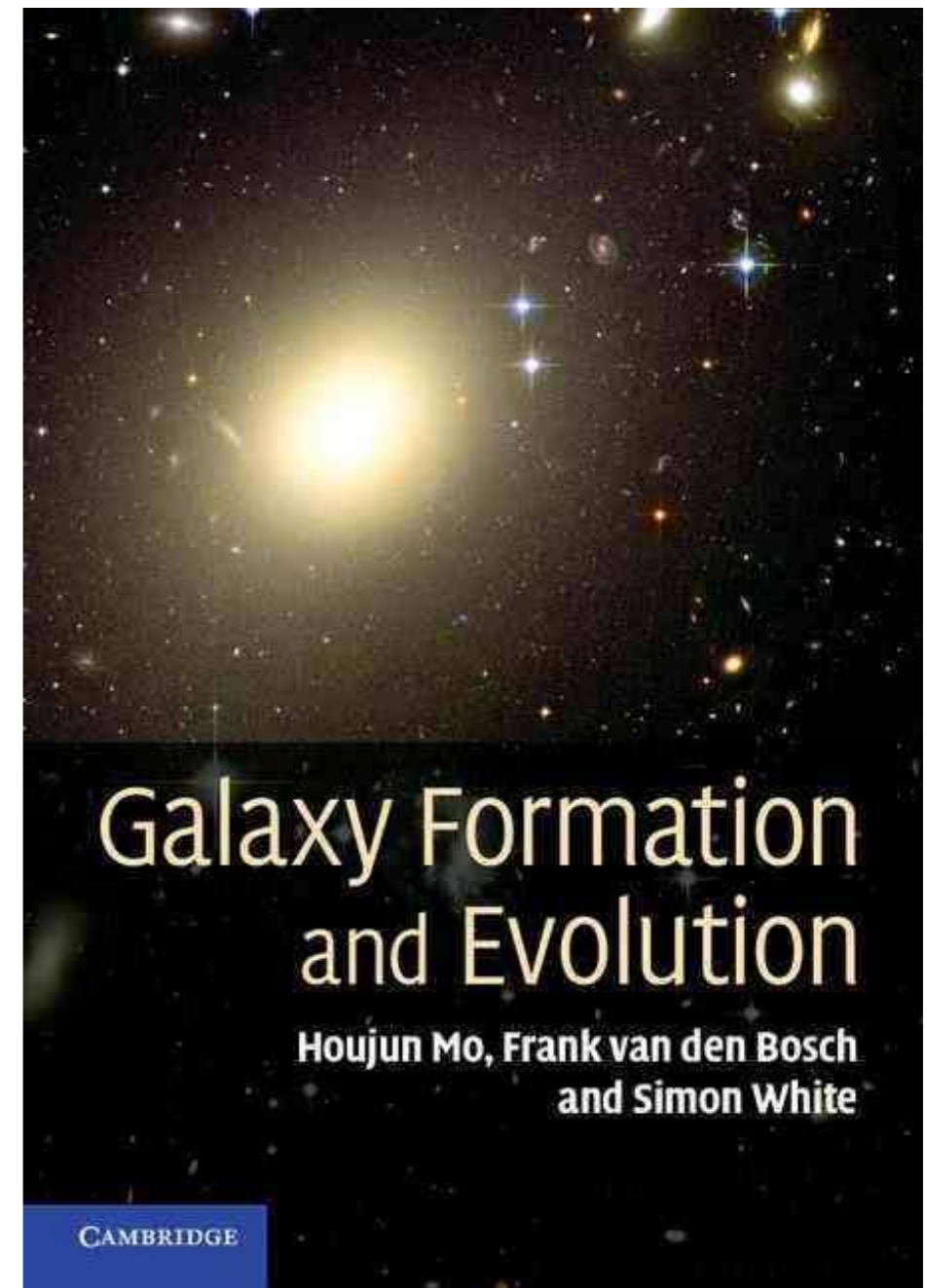
# Additional reading

- Binney & Merrifield, *Galactic Astronomy*, 1998, Princeton University Press
- Will use for galaxy phenomenology and topics related to galaxy evolution / formation



# Additional reading

- Mo, van den Bosch, & White, *Galaxy Formation and Evolution*, 2010, Cambridge University Press
- Will use for formation of dark-matter halos and topics related to galaxy evolution / formation
- Additional readings indicated on the course website



# Code

- Lecture notes contain code examples in Python
- Assignments + final will require some coding as well, preferably done in Python (e.g., jupyter notebook)
- Please try installing all of the pre-requisites ASAP, don't wait until the first assignment

# Code

- Require:

```
name: galdyncourse
channels:
  - conda-forge
dependencies:
  - python=3.8
  - ipython
  - jupyter
  - numpy
  - scipy
  - matplotlib
  - pyqt
  - pandas
  - pip
  - conda-forge::astropy>=2
  - conda-forge::galpy
  - pip:
    - astroquery
```

- See the Notes' Preface for an `environment.yml` file that easily allow you to setup a **conda** environment for this course that contains everything you need
- (note that because some of the code required in the notes is in active development, for some code examples you will need to download galpy's development version)
- The course notes have live code

# Marking scheme

- Assignments: 3 assignments throughout the semester —> total 30%
- Participation —> 20%:
  - Each student needs to send at least one question by the end of the work day the day before the Q&A session
- Presentation: Each student gives a short presentation in week 10 (Nov. 23-27, exact date/time TBD) on a topic in “Galactic Structure and Dynamics”; we’ll discuss possible topics later —> 20% of total
- Take-home final + oral —> 30%

# Assignments

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- **Assignment 1:** TBP on Oct. 1, due Oct 15 at the start of class
- **Assignment 2:** TBP on Oct. 22, due Nov. 5 at the start of class
- **Assignment 3:** TBP on Nov. 5, due Nov. 19 at the start of class

(preliminary) Schedule



# Schedule

- **Week 1:** Class logistics; Introduction to galactic structure; overview of background knowledge.
- **Week 2:** General properties of gravitational potentials; properties and examples of spherical mass distributions; basics of classical mechanics; orbits in spherical potentials.
- **Week 3:** Galaxies as collisionless systems; equilibrium configurations of spherical systems; virial theorem; collisionless Boltzmann equation; spherical Jeans equations; spherical distribution functions; applications: masses of spherical systems.
- **Week 4:** Properties of disk mass distributions; orbits in axisymmetric potentials; dark matter; rotation curves; gas kinematics in the Milky Way.
- **Week 5:** Asymmetric drift; the dynamics of the solar neighborhood; Spheroidal and triaxial mass distributions; orbits in these mass distributions; surfaces of section; chaos; Schwarzschild modeling.
- **Week 6:** Numerical methods;  $N$ -body modeling.
- **Week 7:** Formation and evolution of dark matter halos; violent relaxation; phase-mixing.
- **Week 8:** Chemical evolution of galaxies; age–abundance relations in the solar neighborhood; stellar population synthesis.
- **Week 9:** Internal structure of elliptical galaxies; supermassive central black holes; stability of stellar systems; bars; spiral arms.
- **Week 10:** Student presentations.
- **Week 11:** Mergers and dynamical friction; tides.
- **Week 12:** Review.

## What's going on this week? What's due?

Week	Dates	Topic	Due on Thu?
1	Sep 14 – Sep 18	Introduction to galactic structure	
2	Sep 21 – Sep 25	Gravitation, classical mechanics, spherical orbits	
3	Sep 28 – Oct 02	Dynamical equilibria, masses of spherical systems	
4	Oct 05 – Oct 09	Galactic disks, galactic rotation	
5	Oct 12 – Oct 16	Disk equilibria, spheroidal mass distributions	Assignment 1
6	Oct 19 – Oct 23	Numerical methods	Presentation topic
7	Oct 26 – Oct 30	Dark matter halos	
8	Nov 02 – Nov 06	Chemical evolution	Assignment 2
	Nov 09 – Nov 13	<b>Reading week, no class</b>	
9	Nov 16 – Nov 20	Elliptical galaxies, stability	Assignment 3
10	Nov 23 – Nov 27	<b>Presentations</b>	
11	Nov 20 – Dec 04	Mergers and dynamical friction	
12	Dec 07 – Dec 11	Review	

Pop Quiz!

What is the diameter of the  
Milky Way disk?

- A. 3 kpc
- B. 10 kpc
- C. 30 kpc
- D. 100 kpc

# What is the diameter of the Milky Way disk?

A. 3 kpc

B. 10 kpc

C. 30 kpc

D. 100 kpc

How thick is the Milky Way  
disk?

A. 100 pc

B. 600 pc

C. 2 kpc

D. 20 kpc

# How thick is the Milky Way disk?

A. 100 pc

B. 600 pc

C. 2 kpc

D. 20 kpc

How many stars does the  
Milky Way contain?

A.  $10^5$

B.  $10^7$

C.  $10^{11}$

D.  $10^{13}$



# How many stars does the Milky Way contain?

A.  $10^5$

B.  $10^7$

C.  $10^{11}$

D.  $10^{13}$

What is the ratio of (dark matter) / (stellar matter) in total in the Milky Way?

A. 0.3

B. 1

C. 3

D. 15

What is the ratio of (dark matter) / (stellar matter) in total in the Milky Way?

A. 0.3

B. 1

C. 3

D. 15

What is the orbital period of the Sun around the Galactic center?

- A. 1 Gyr
- B. 100 Myr
- C. 50 Myr
- D. 250 Myr

What is the orbital period of the Sun around the Galactic center?

A. 1 Gyr

B. 100 Myr

C. 50 Myr

D. 250 Myr